

# Инструкция по изменению данных в Thomson Reuters Web of Science

К.А. Постнов, ГАИШ МГУ

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Готово

both the turn-ons and the variations in pulse profiles. Is there really free precession in the NS? (Staubert et al., Proc. INTEGRAL Conf. Dublin).

27  
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## Шаг 2.

- База данных выдаст список Ваших работ в заказанный диапазон лет (по умолчанию – с 1898 г по н.в.)

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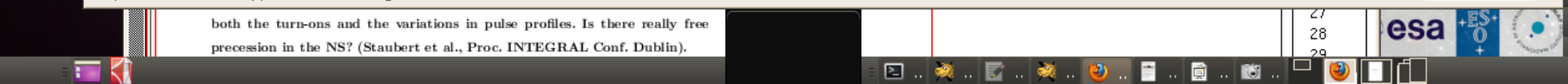
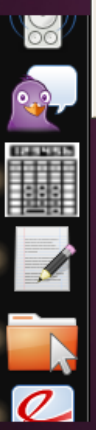
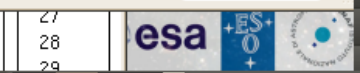
**Document Types**

**Authors**

- Authors - Chinese
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- Title: **Quasi-periodic flares in EXO2030+375 observed with INTEGRAL**  
 Author(s): Klochkov D.; Ferrigno C.; Santangelo A.; et al.  
 Source: ASTRONOMY & ASTROPHYSICS Volume: 536 Article Number: L8 DOI: 10.1051/0004-6361/201118185 Published: DEC 2011  
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 Source: ASTRONOMY & ASTROPHYSICAL JOURNAL LETTERS Volume: 742 Issue: 1 Article Number: L11 DOI: 10.1088/2041-8205/742/1/L11  
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 Author(s): Revnivtsev M.; Postnov K.; Kuranov A.; et al.  
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 Author(s): Pshirkov M. S.; Postnov K. A.  
 Source: ASTRONOMY AND SPACE SCIENCE Volume: 330 Issue: 1 Pages: 13-18 DOI: 10.1007/s10509-010-0395-x Published: NOV 2010  
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both the turn-ons and the variations in pulse profiles. Is there really free precession in the NS? (Staubert et al., Proc. INTEGRAL Conf. Dublin).



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- Нажмите на название работы

## THE 5 hr PULSE PERIOD AND BROADBAND SPECTRUM OF THE SYMBIOTIC X-RAY BINARY 3A 1954+319

**Author(s):** Marcu, DM (Marcu, Diana M.)<sup>1,2,3</sup>; Furst, F (Fuerst, Felix)<sup>4,5</sup>; Pottschmidt, K (Pottschmidt, Katja)<sup>1,2,3</sup>; Grinberg, V (Grinberg, Victoria)<sup>4,5</sup>; Muller, S (Mueller, Sebastian)<sup>4,5</sup>; Wilms, J (Wilms, Joern)<sup>4,5</sup>; Postnov, KA (Postnov, Konstantin A.)<sup>6</sup>; Corbet, RHD (Corbet, Robin H. D.)<sup>1,2,3</sup>; Markwardt, CB (Markwardt, Craig B.)<sup>1</sup>; Bel, MC (Cadolle Bel, Marion)<sup>7</sup>

**Source:** ASTROPHYSICAL JOURNAL LETTERS **Volume:** 742 **Issue:** 1 **Article Number:** L11 **DOI:** 10.1088/2041-8205/742/1/L11  
**Published:** NOV 20 2011

**Times Cited:** 0 (from Web of Science)

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**Abstract:** We present an analysis of the highly variable accreting X-ray pulsar 3A 1954 + 319 using 2005-2009 monitoring data obtained with INTEGRAL and Swift. This considerably extends the pulse period history and covers flaring episodes in 2005 and 2008. In 2006 the source was identified as one of only a few known symbiotic X-ray binaries (SyXBs), i.e., systems composed of a neutron star accreting from the inhomogeneous medium around an M-giant star. The extremely long pulse period of similar to 5.3 hr is directly visible in the 2008 INTEGRAL-ISGR1 outburst light curve. The pulse profile is double peaked and generally not significantly energy dependent although there is an indication of possible softening during the main pulse. During the outburst a strong spin-up of  $-1.8 \times 10^{-4}$  hr hr<sup>-1</sup> occurred. Between 2005 and 2008 a long-term spin-down trend of  $2.1 \times 10^{-5}$  hr hr<sup>-1</sup> was observed for the first time for this source. The 3-80 keV pulse peak spectrum of 3A 1954 + 319 during the 2008 flare could be well described by a thermal Comptonization model. We interpret the results within the framework of a recently developed quasi-spherical accretion model for SyXBs.

**Accession Number:** WOS:000296763600011

**Document Type:** Article

**Language:** English

**Author Keywords:** binaries: symbiotic; stars: individual (3A 1954+319); stars: neutron; X-rays: binaries

**KeyWords Plus:** LONG-PERIOD; CATALOG; 4U-1954+319; CYGNUS-X-1; 4U-1700+24; MISSION; GX-1+4; GIANTS; STARS; BAT

**Reprint Address:** Marcu, DM (reprint author), NASA, Goddard Space Flight Ctr, Astrophys Sci Div, Code 661, Greenbelt, MD 20771 USA

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6. Sternberg Astron Inst, Moscow 119992, Russia
7. European Space Agcy, European Space Astron Ctr, Madrid 28692, Spain

### Funding:

Funding Agency	Grant Number
NASA	NNX08AE84G NNX08AY24G NNX09AT28G
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